

High-Accuracy Mass Measurements with SMILETRAP

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Introduction to SMILETRAP

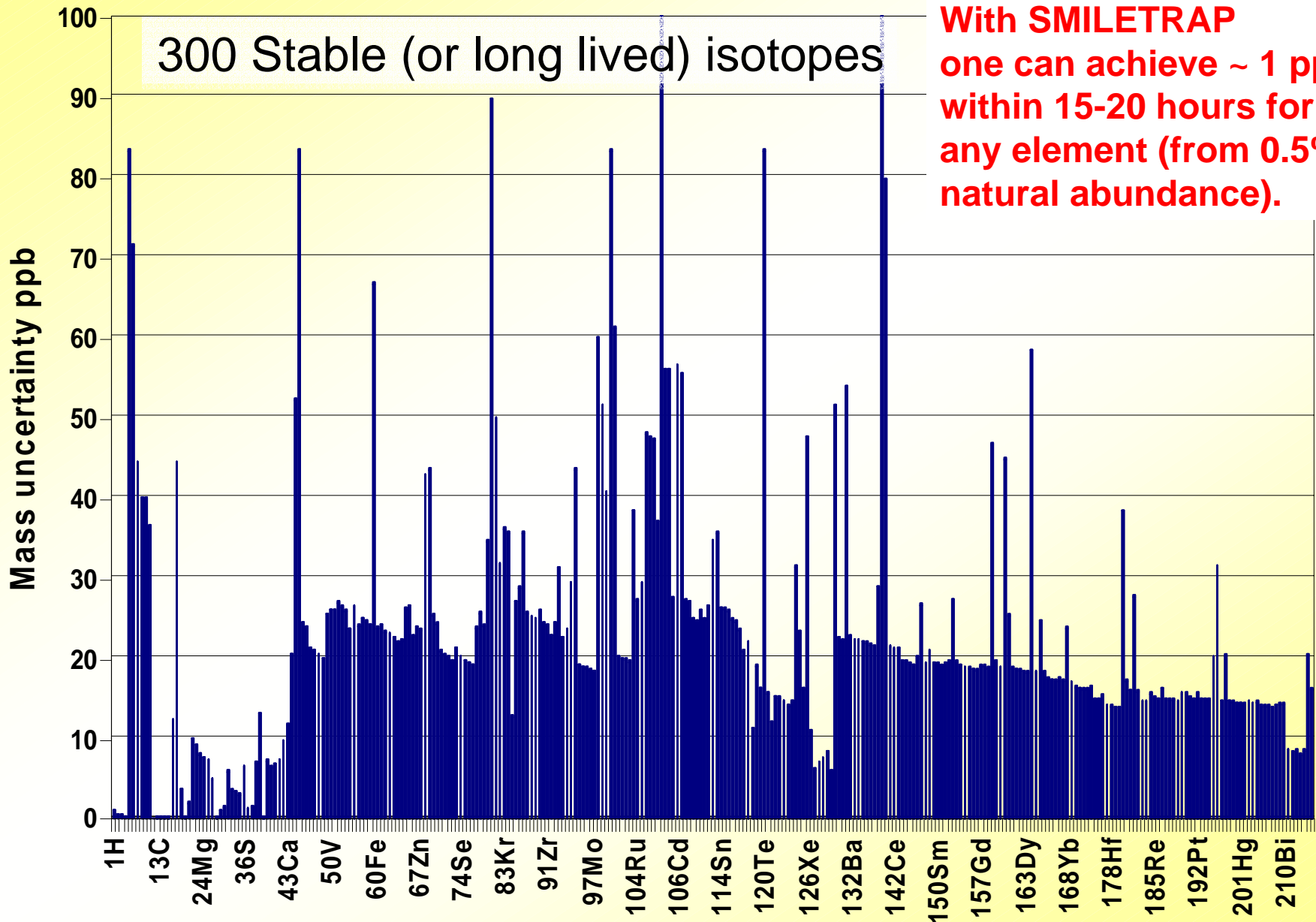
- **Objective** : High-precision (ppb) atomic mass spectrometry using highly charged ions
- **Principle** : Measurement of the cyclotron frequency of an ion trapped in a homogeneous magnetic field :

$$\nu_c = \frac{1}{2\pi} \frac{qeB}{m} \qquad \frac{m}{\Delta m} = \frac{\nu_c}{\Delta \nu_c}$$

→ using HCl increases the precision linearly

- **The experiment** :
 - A CHORDIS source to produce high intensity beam of mono-isotopic singly charged ions
 - An EBIS to produce HCl, $\sim U^{65+}$ (Neon like ions)
 - A PENNING trap

Current knowledge of atomic masses from most recent compilation (1997)



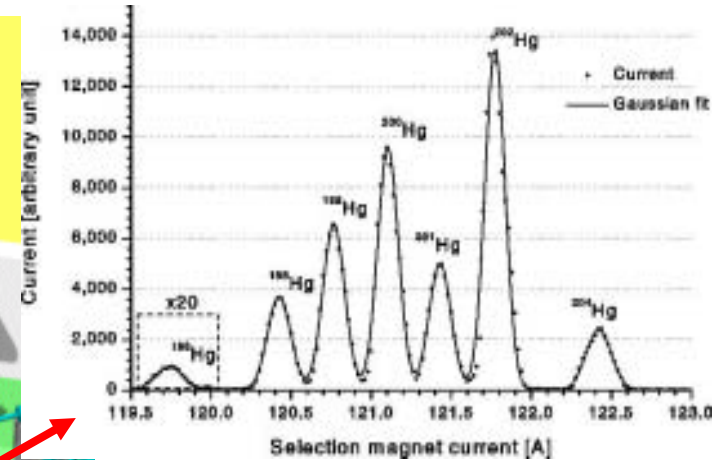
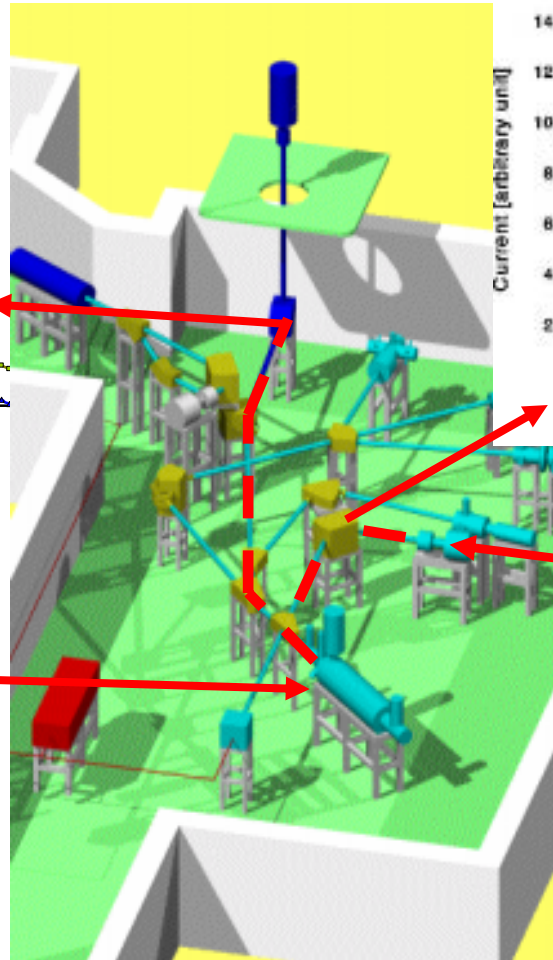
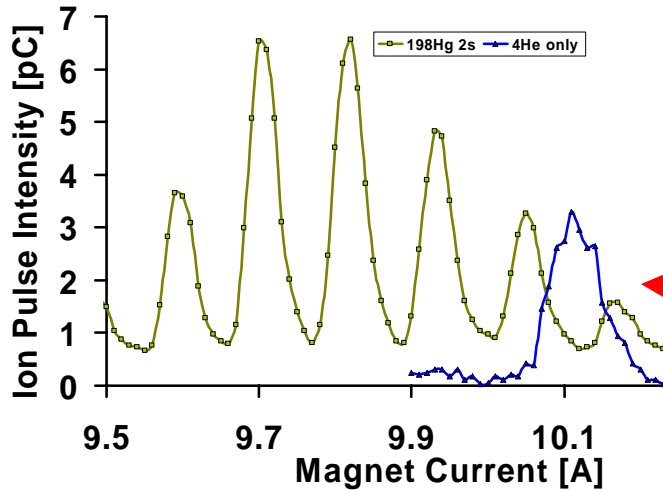
**With SMILETRAP
one can achieve ~ 1 ppb
within 15-20 hours for
any element (from 0.5%
natural abundance).**

Where does the mass of an atom or ion matter?

- ^{28}Si for atomically defined kilogram mass standard
 - ^{76}Ge for constraints on neutrino-less double beta decay
 - ^{133}Cs , for accurate determination of the fine structure constant α
 - ^{24}Mg and ^{26}Mg for bound-electron g factor determination in hydrogen-like ions
 - $^{198-204}\text{Hg}$ to solve the “mercury problem” in Audi/Wapstras mass table
 - Determination of the electron binding energies in heavy atoms
-

... a relative mass accuracy of $\Delta m/m = 10^{-9}-10^{-10}$ is required

The CHORDIS/CRYISIS ion sources:



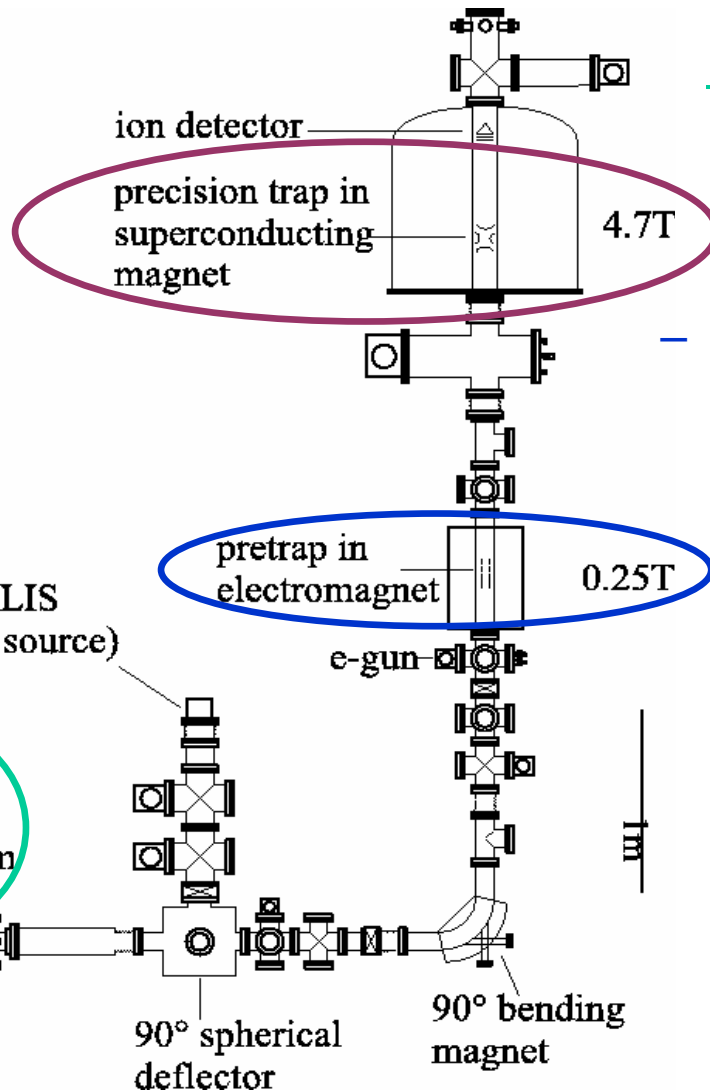
CRYISIS

Production of highly charged ions by electron bombardment

CHORDIS

Production of singly charged isotope separated ions

SMILETRAP setup and Ion budget



– EBIS

- Out from CRYISIS : 10^8 Ge ions
- 25% on Ge^{22+} : 2.5×10^7 ions

– PreTrap

- After magnet and retardation : 10^6 ions
- 1/250 beam captured : < 4000
 - pretrap length/beam length
 - 2 V deep trap / 7 V energy spread

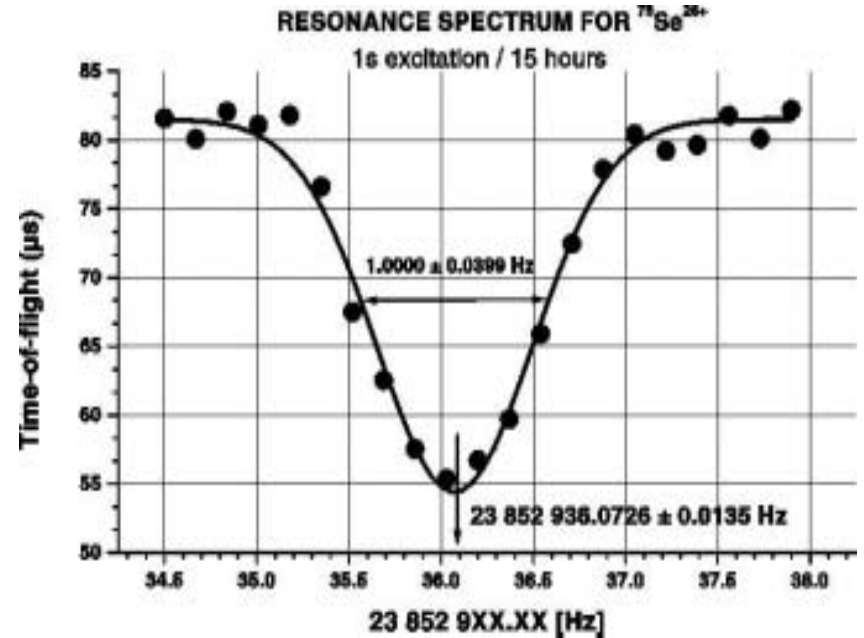
– Precision Trap

- After injection aperture : 150 ions
- Captured : 50 ions
- After energy selection : 1-4 ions
 - 50 mV / 2 V

How to determine the mass of an atom?

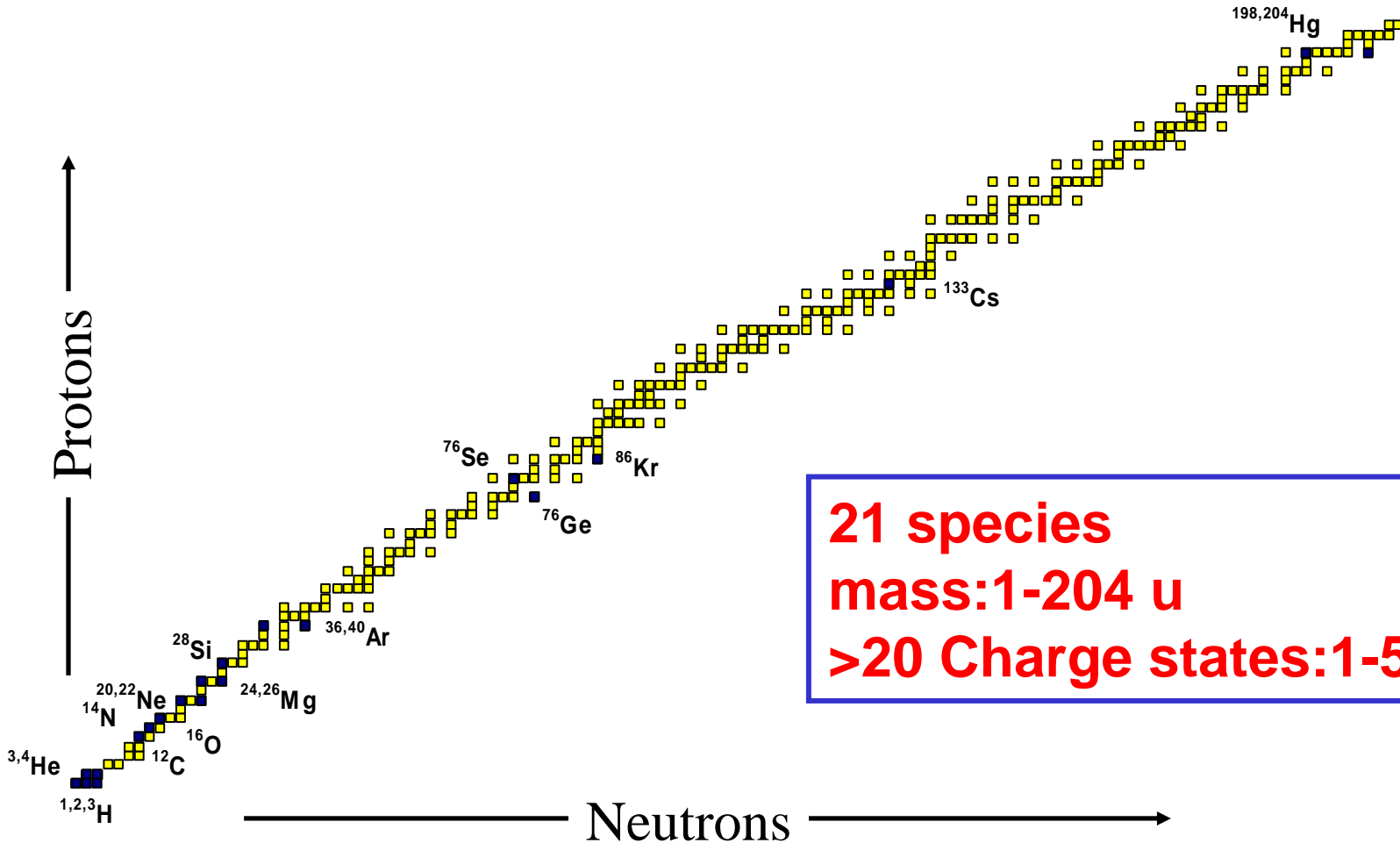
$$V_c = \frac{1}{2\pi} \frac{qeB}{m} \quad \Delta V_c [\text{Hz}] \approx \frac{1.0}{T_{exc} [s]}$$

$$\frac{m}{\Delta m} = \frac{V_c}{\Delta V_c}$$



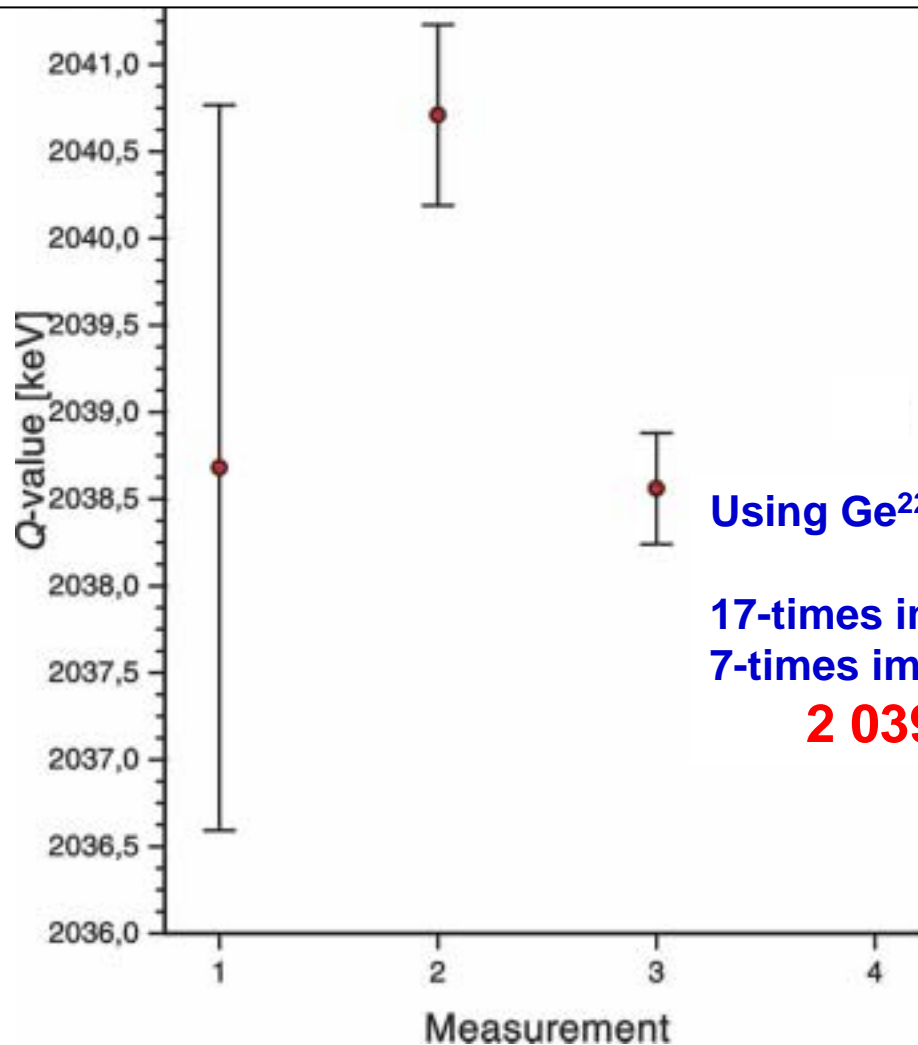
$$R = \frac{V_c}{V_{cREF}} = \frac{qm_{REF}}{q_{REF}m} \implies M_A = \frac{1}{R} \frac{q}{q_{REF}} m_{REF} + qm_e - E_B(A^{q+})$$

Masses measured at SMILETRAP 1997-2002



^{76}Ge for constraints on neutrino-less double beta decay

Question: Is there a ν -less double β -decay?

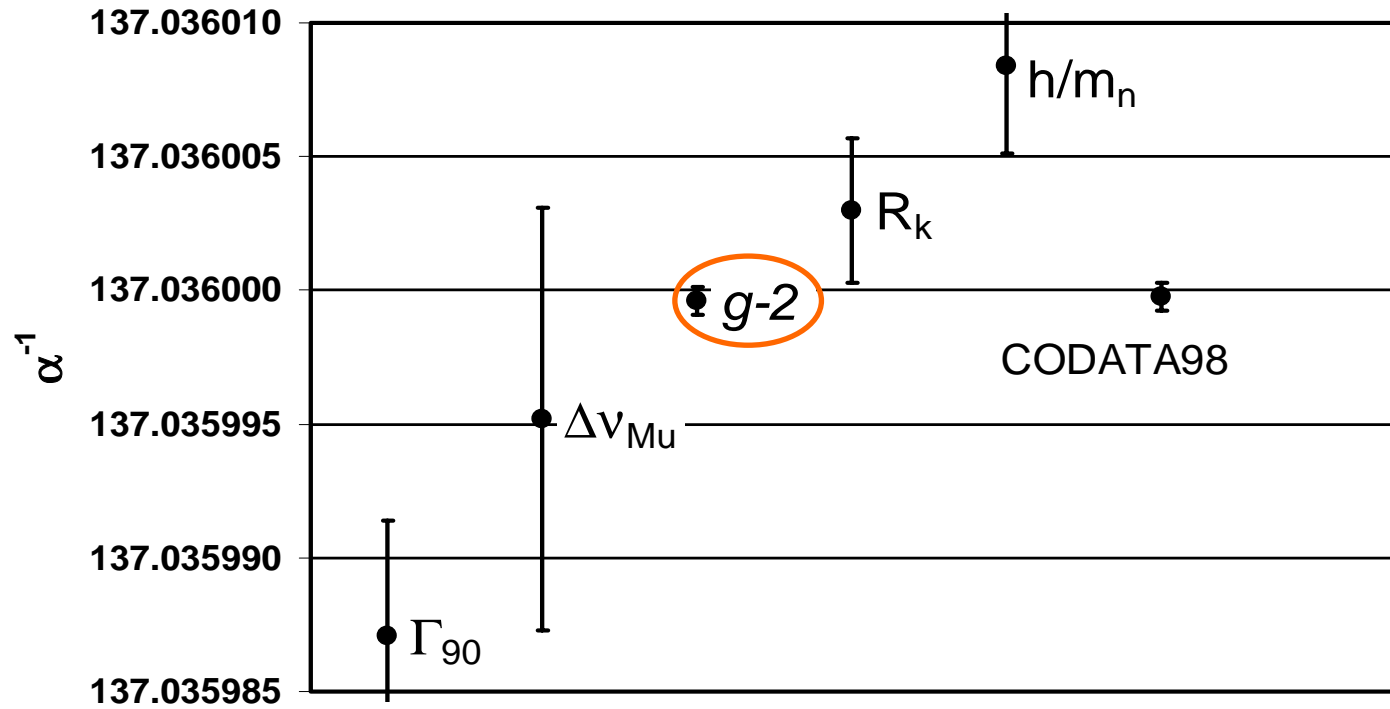


Using $\text{Ge}^{22,23+}$ and $\text{Se}^{24,25+}$ ions:

17-times improvement in both masses
7-times improvement in the Q-value

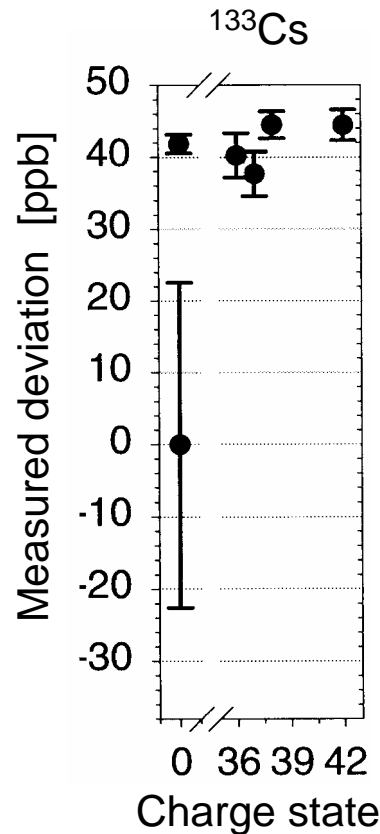
2 039.006(50) keV

Present status of the fine structure Constant α



^{133}Cs for accurate determination of the fine structure constant α

$$\alpha^2 = \left(\frac{2R_\infty}{c}\right) \left(\frac{h}{m_e}\right) = \left(\frac{2R_\infty}{c}\right) \left(\frac{h}{m_{\text{Cs}}}\right) \left(\frac{m_{\text{Cs}}}{m_p}\right) \left(\frac{m_p}{m_e}\right)$$



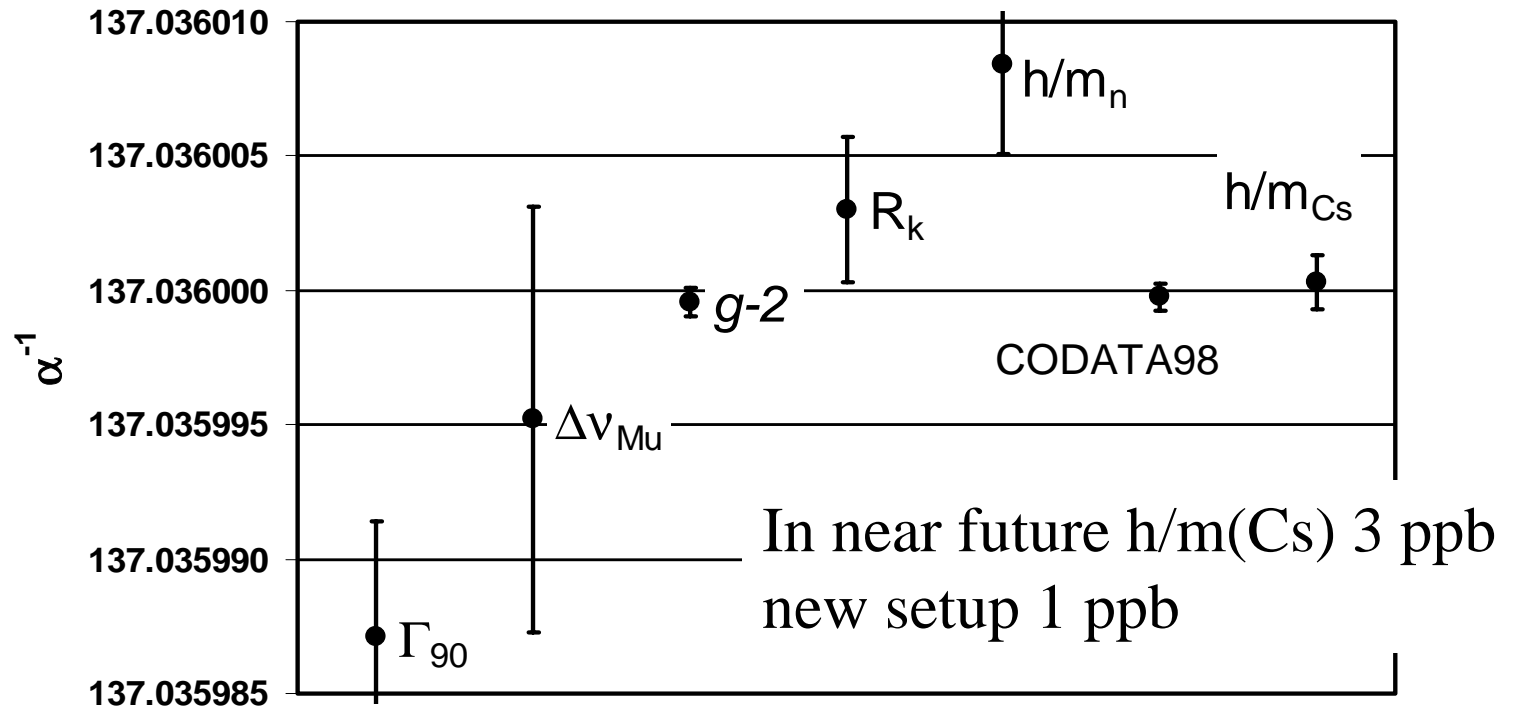
2 ppb (0.4ppb by M. P. Bradley et al.,)

^{133}Cs for accurate determination of the fine structure constant α

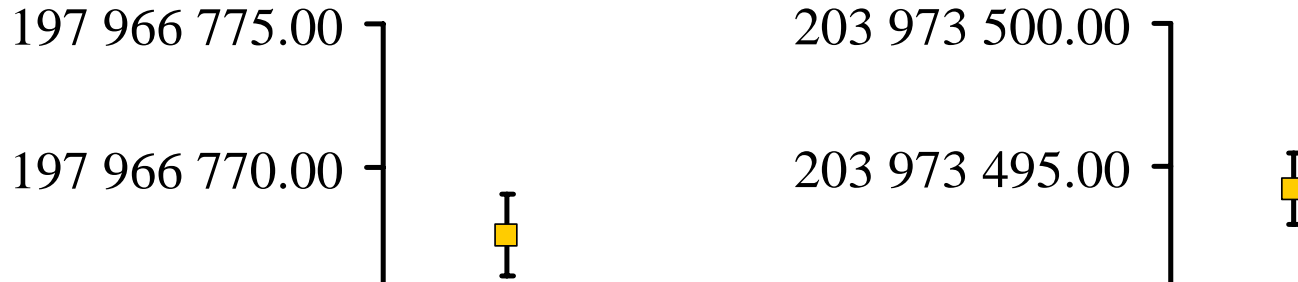
$$\alpha^2 = \left(\frac{2R_\infty}{c} \right) \left(\frac{h}{m_e} \right) = \left(\frac{2R_\infty}{c} \right) \left(\frac{h}{m_{\text{Cs}}} \right) \left(\frac{m_{\text{Cs}}}{m_p} \right) \left(\frac{m_p}{m_e} \right)$$

Annotations for the equation above:

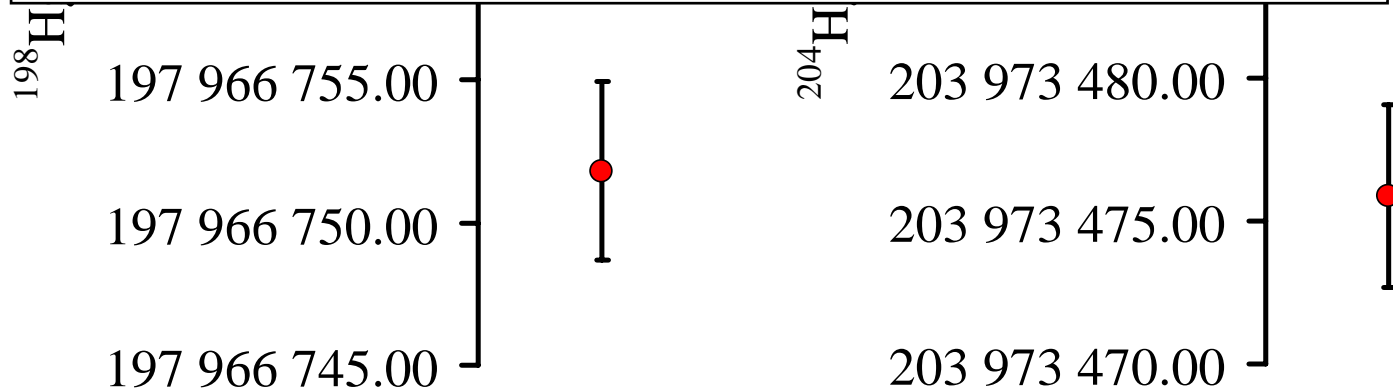
- $2R_\infty$: $< 2 \cdot 10^{-11}$
- h : 7.3 ppb
- m_{Cs} : 2 ppb (0.4 ppb by M. P. Bradley et al.)
- m_p : 2.2 ppb
- m_e : 0.7 ppb from g-factor, C^{5+}



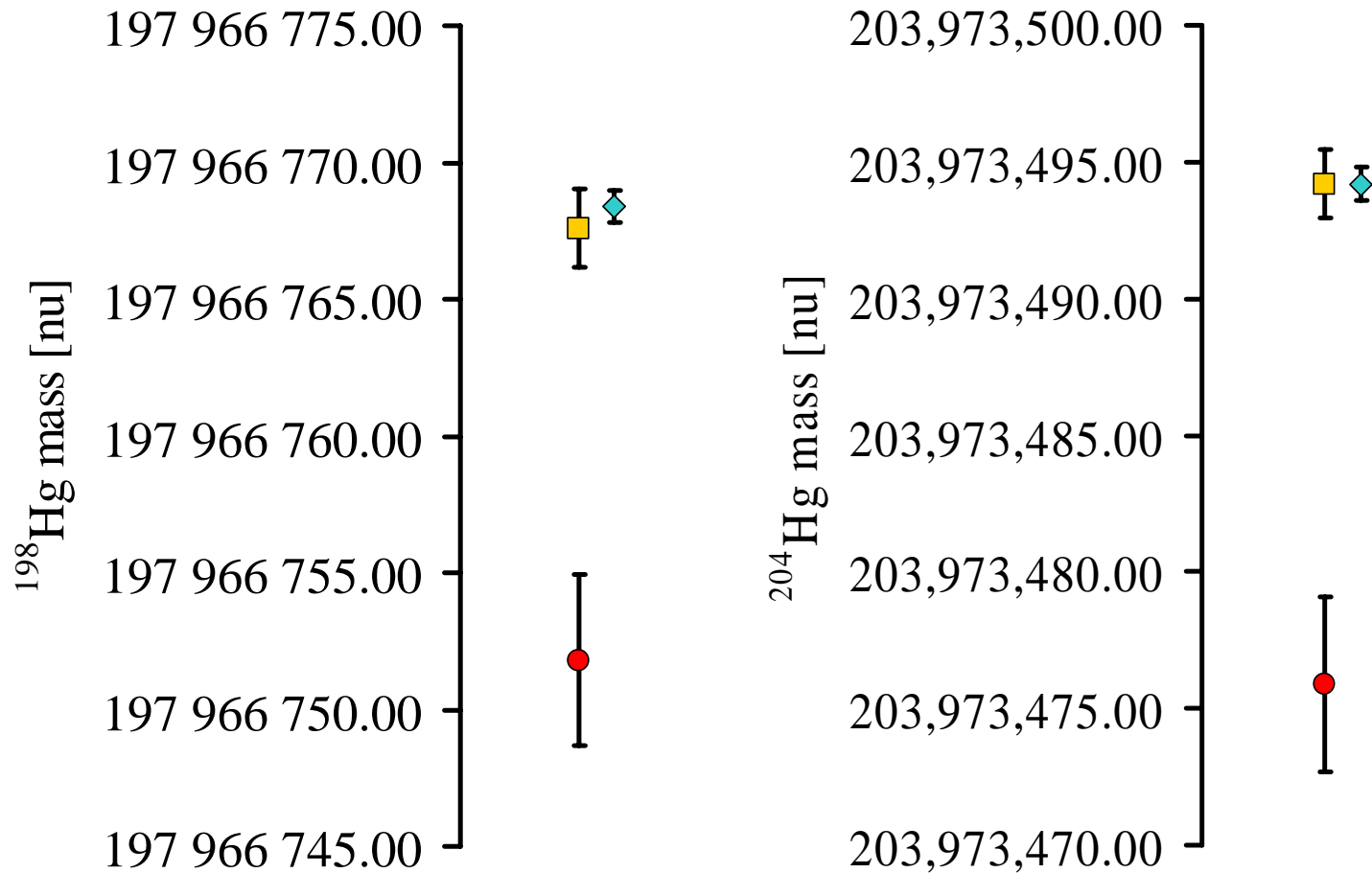
The “mercury problem” in Audi and Wapstras mass table



We do not feel happy about the situation and think that a re-measurement of the Hg masses combined with that of lighter elements ($Z=73-77$), is the single most desirable experiment concerning masses near the line of β -stability.
Audi & Wapstra 1993.



Is the “mercury problem” solved?

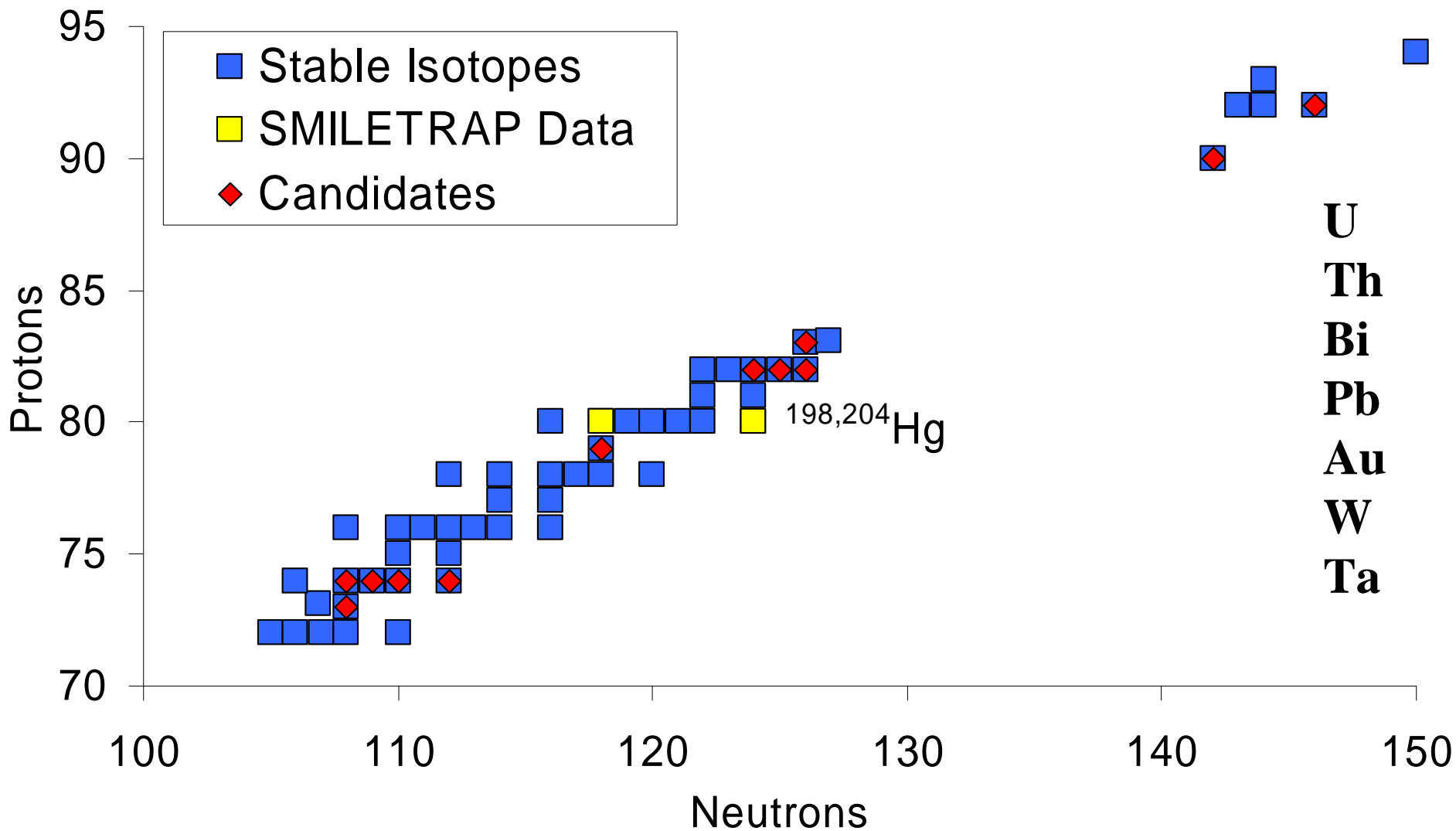


Manitoba results confirmed



$\Delta m/m = 2 \cdot 10^{-9}$

Important candidates to solve the “mercury problem”



Determination of electron binding energies

New tests on theory?

Examples:

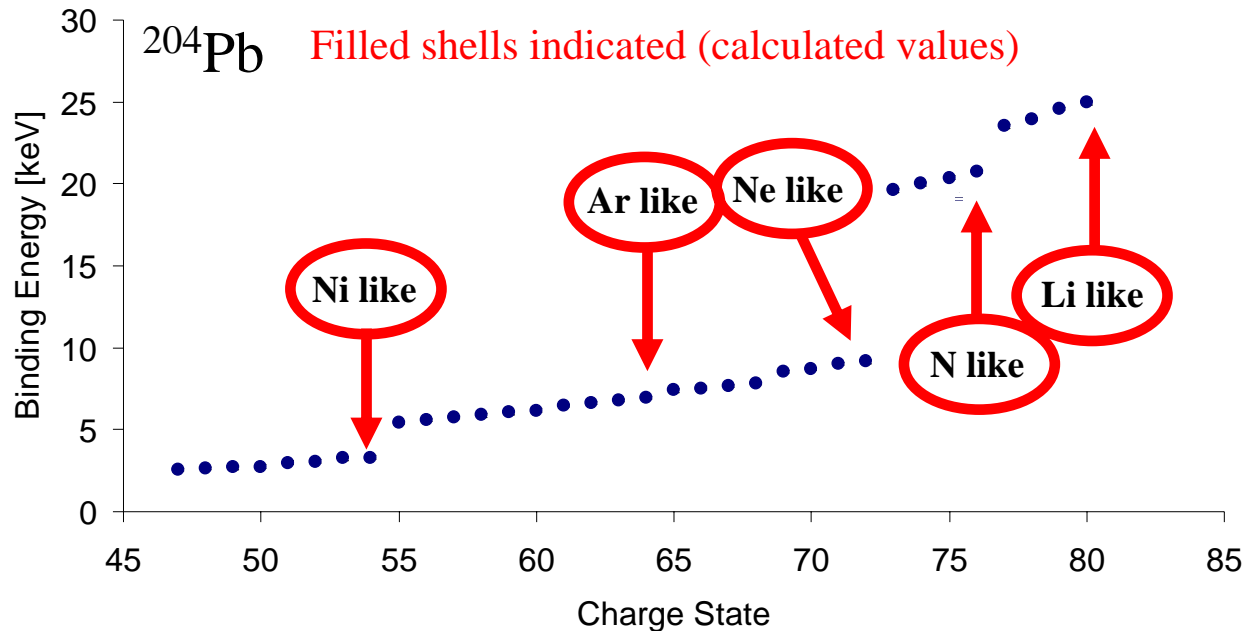
$$\text{H}2^+ \rightarrow E_B = -15.4 \text{ eV} = 8.3 \text{ ppb}$$

$$^{208}\text{Pb}^{50+} \rightarrow E_B = -50.372 \text{ keV} = 260 \text{ ppb}$$

$$^{208}\text{Pb}^{72+} \rightarrow E_B = -172.177 \text{ keV} = 889 \text{ ppb}$$

Electron binding energies (E_B) are needed to calculate the mass of a neutral atom:

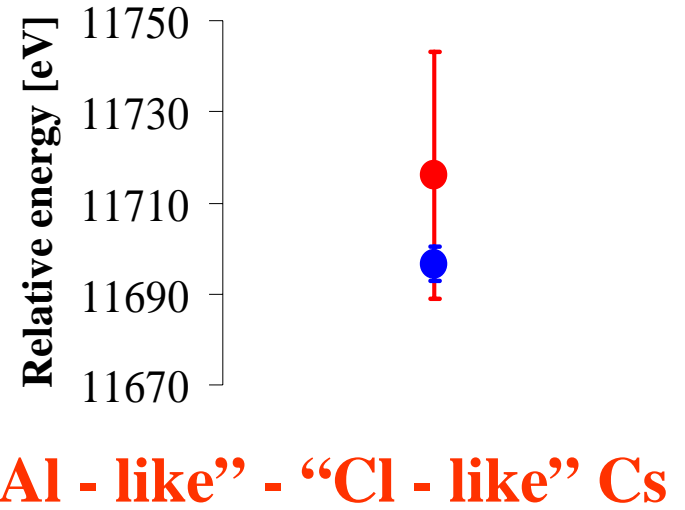
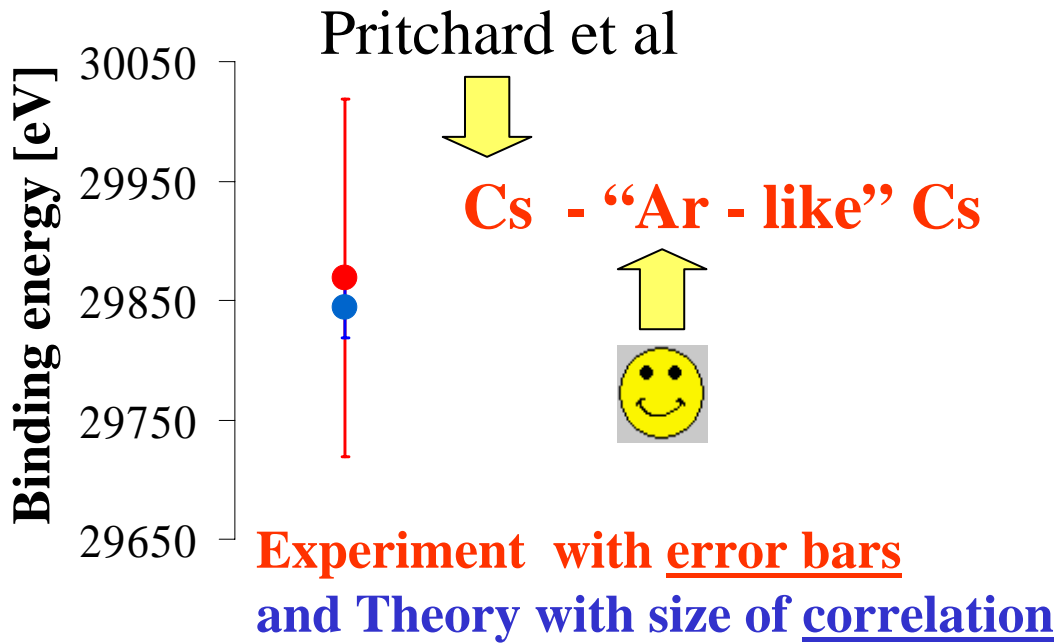
$$M_A = \frac{1}{R} \frac{q}{q_{ref}} - qm_e - E_B(A^{q+})$$



Examples where calculations are used (tested)

Measure :Hg⁵²⁺ (..3d¹⁰), Add: 52 m_e and E_B= -56 377(50) eV

Binding energy 56377.66 eV uncertainty ~50 eV = 0.5 ppb
 Scofield's values deviates 2 ppb SMILETRAPs mass=2 ppb



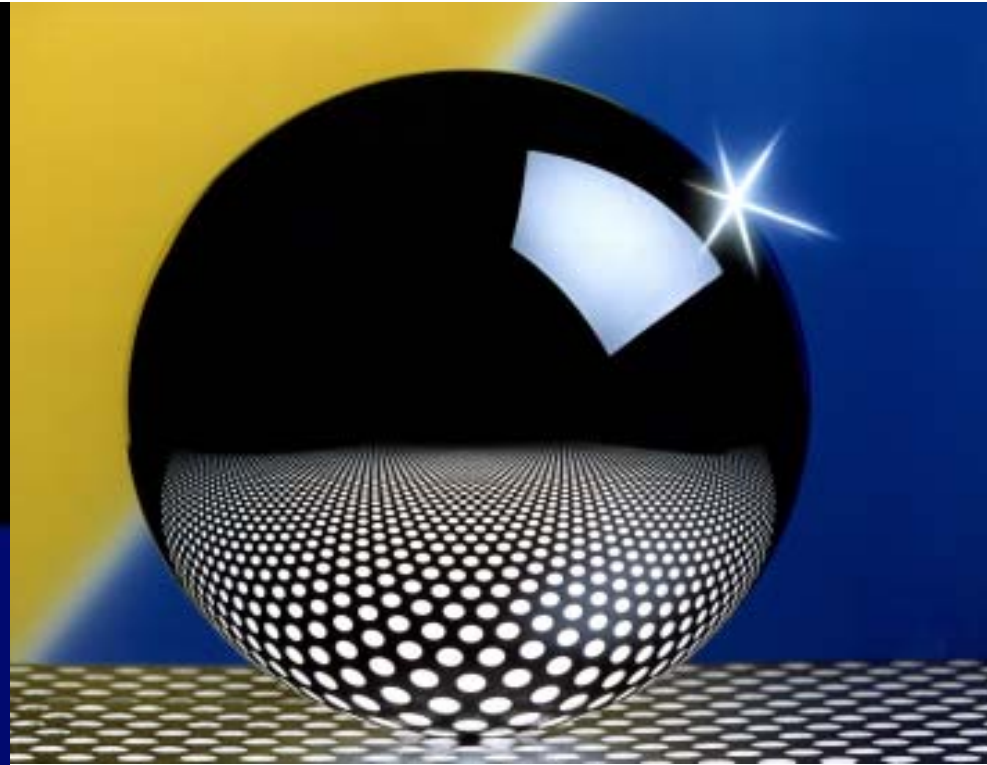
Conclusion

- **The mass is a fundamental property that is of importance in many fields in physics**
- **With SMILETRAP almost any stable mass can be measured with an uncertainty <1 ppb**
In a few case an accuracy close to 10^{-10} have been reached
- **Next mass determinations :**
 - Some mass(es) in the mercury region such as Ta, Pb, ^{232}Th or $^{235,238}\text{U}$ to solve the mass table dilemma
 - Test of QED in strong fields, ^{40}Ca for g-factor measurement
- **High precision measurements of e⁻ binding energies :**
 - Almost any heavy and highly charged atom, mass >100 u
 - 0.2 ppb (~20 eV for mass 100) uncertainty in energy is within reach
 - Test of theoretical calculations

$$M_A = \frac{1}{R} \frac{q}{q_{REF}} m_{REF} + qm_{e^-} - E_B(A^{q+})$$

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- ^{28}Si for New Definition of Kilogram Mass Standard



Avogadro Constant

$$N_A = \frac{M(\text{Si}) \cdot \rho}{V_0 \cdot n}$$

$M(\text{Si})$ molar mass of silicon

ρ the density of the silicon body

V_0 the volume and n the number of atoms of a unit cell

^{28}Si 1×10^{-10} direct measurement by MIT
 ^{29}Si and ^{30}Si 1×10^{-9} indirectly determined

SMILETRAP: ^{28}Si $\sim 3 \times 10^{-10}$ using HCl

^{29}Si and ^{30}Si should also be measured or a pure ^{28}Si crystal can be used