

# Masses of atomic nuclei far from stability

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Introduction: Models for nuclear masses.

Symmetry energy in nuclear matter.

Chaotic mass component?

Mass predictions from pattern recognition.

# ECT\* doctoral training programme

- Title: “Nuclear structure and reactions” (8<sup>th</sup> March-8<sup>th</sup> June 2007, for PhD students).
- Lecture series on shell model, mean-field models, nuclear astrophysics, symmetries in nuclei, reaction theory, exotic nuclei, open quantum systems, fundamental interactions.
- Workshops related to these topics.
- Please:
  - Encourage students to apply;
  - Submit workshop proposals to ECT\*.

# Nuclear mass formulas

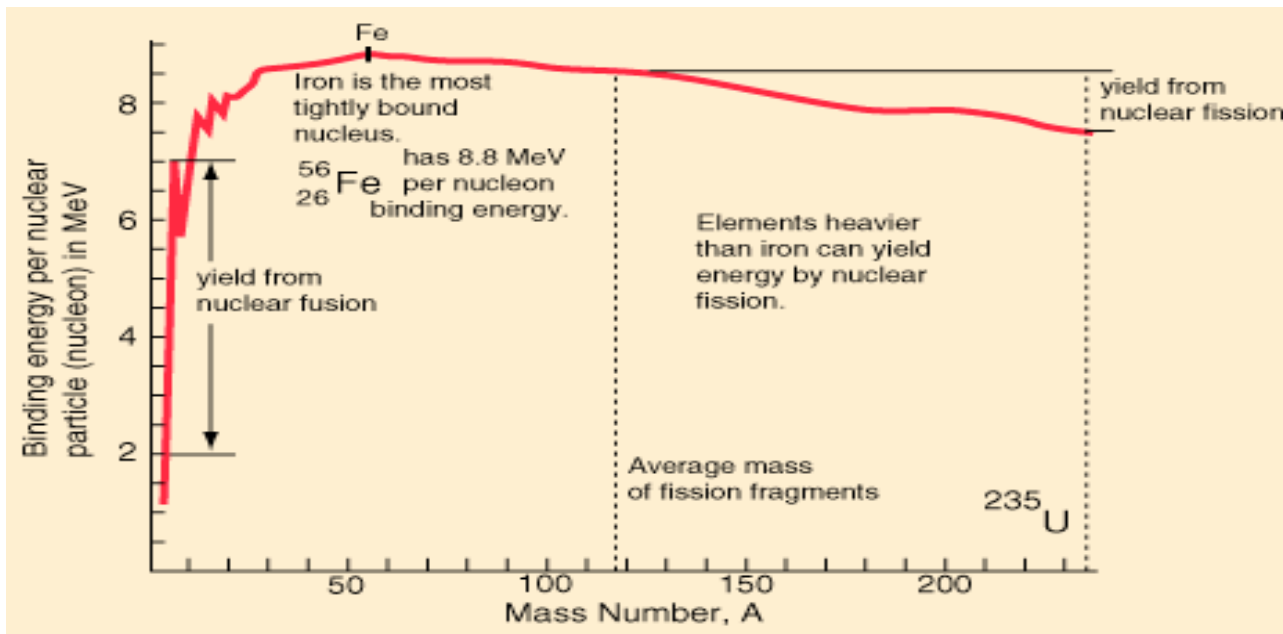
- Global mass formulas:
  - Liquid-drop model (LDM): von Weizsäcker.
  - Macroscopic models with microscopic corrections: FRDM, ...
  - Microscopic models: HFB $n$ , RMF, DZ, ...
- Local mass formulas:
  - Extrapolations by Wapstra & Audi.
  - IMME, Garvey-Kelson (GK) relations, mass formula of Liran-Zeldes, neural networks, ...

D. Lunney *et al.*, Rev. Mod. Phys. **75** (2003) 1021  
K. Blaum, Phys. Reports **425** (2006) 1.

# Liquid-drop mass formula

- Binding energy of an atomic nucleus:

$$B(N, Z) = a_{\text{vol}} A - a_{\text{sur}} A^{2/3} - a_{\text{cou}} \frac{Z(Z-1)}{A^{1/3}} - a_{\text{sym}} \frac{(N-Z)^2}{A} + a_{\text{pai}} \frac{\delta(N, Z)}{A^{1/2}}$$



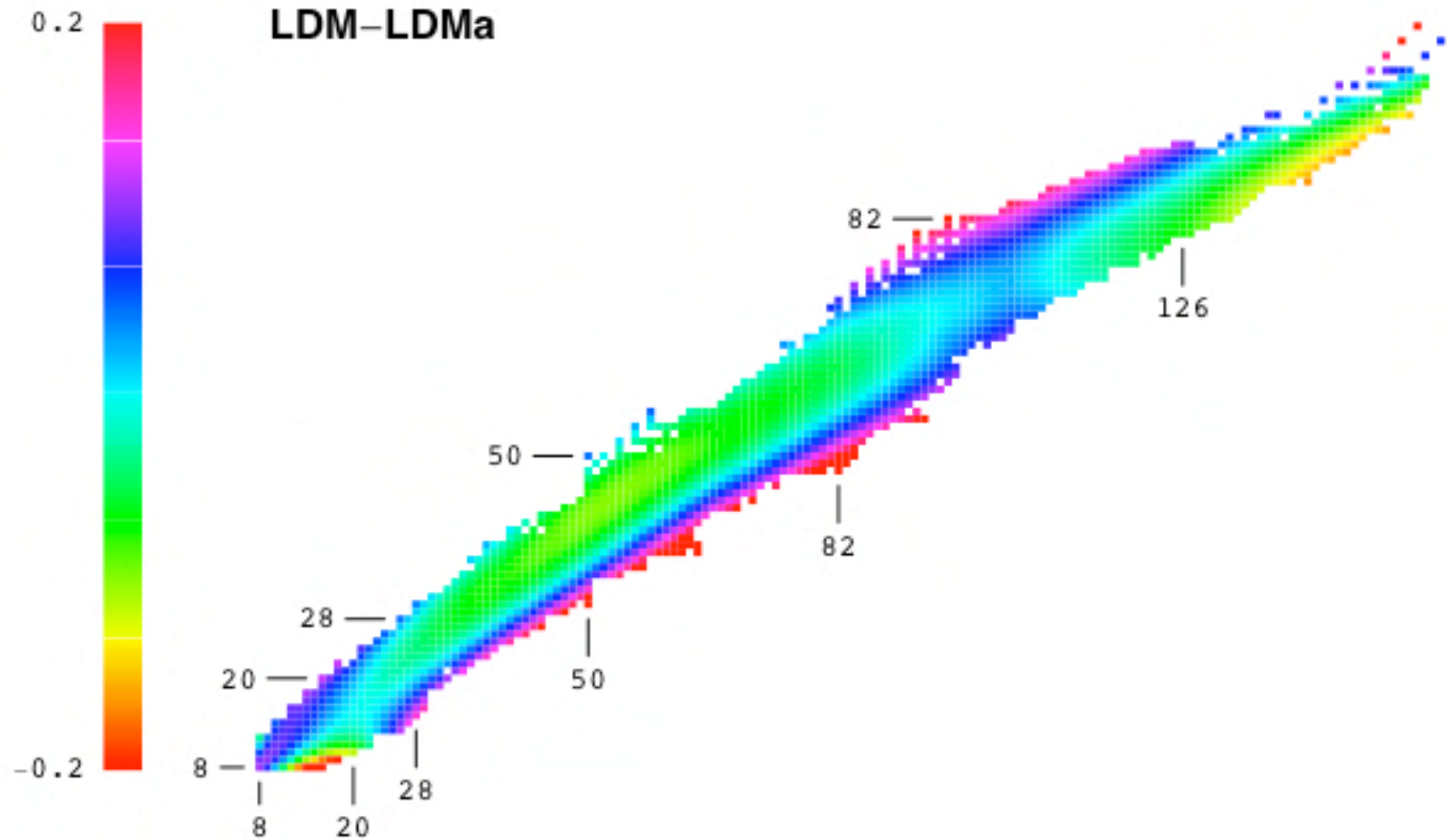
# Deficiencies of Weizsäcker formula

- Consistency of the Weizsäcker mass formula requires a *surface-symmetry* term. Derivation relies on the thermodynamics of an asymmetric two-component system:

$$\frac{S_v}{1 + S_v A^{-1/3} / S_s} \frac{(N - Z)^2}{A} \approx -a_{\text{vsym}} \frac{(N - Z)^2}{4A} + a_{\text{ssym}} \frac{(N - Z)^2}{4A^{4/3}}$$

W.D. Myers & W.J. Swiatecki, *Ann. Phys.* **55** (1969) 395  
A. Bohr & B.R. Mottelson, *Nuclear Structure II* (1975)  
A.W. Steiner *et al.*, *Phys. Reports* **411** (2005) 325  
P. Danielewicz, *Nucl. Phys. A* **727** (2003) 233

# LDM versus LDMa



# Symmetry energy

- Energy per particle in nuclear matter:

$$E(\rho, x) = E\left(\rho, x = \frac{1}{2}\right) + S(\rho)(1 - 2x)^2, \quad x = Z/A$$

- Symmetry energy  $S(\rho)$  is density dependent:

$$S(\rho) = a_4 + p_0(\rho - \rho_0) + \Delta K(\rho - \rho_0)^2$$

- In Thomas-Fermi approximation ( $r=0$ ):

$$\frac{S_v}{S_s} = \frac{3}{R\rho_0} \int dr \rho(r) \left( \frac{S(\rho_0)}{S(\rho)} - 1 \right)$$

# Quantal effects & Wigner cusp

- The  $(N-Z)^2$  dependence of the symmetry term arises in a macroscopic approximation.
- Quantal theories gives rise to
  - $T(T+1)$ : isospin SU(2).
  - $T(T+4)$ : supermultiplet SU(4).
- This suggests a generalization of the form  $T(T+r)$ , with  $r$  a parameter.

N. Zeldes, Phys. Lett. B **429** (1998) 20

J. Jänecke & T.W. O'Donnell, Phys. Lett. B **605** (2005) 87



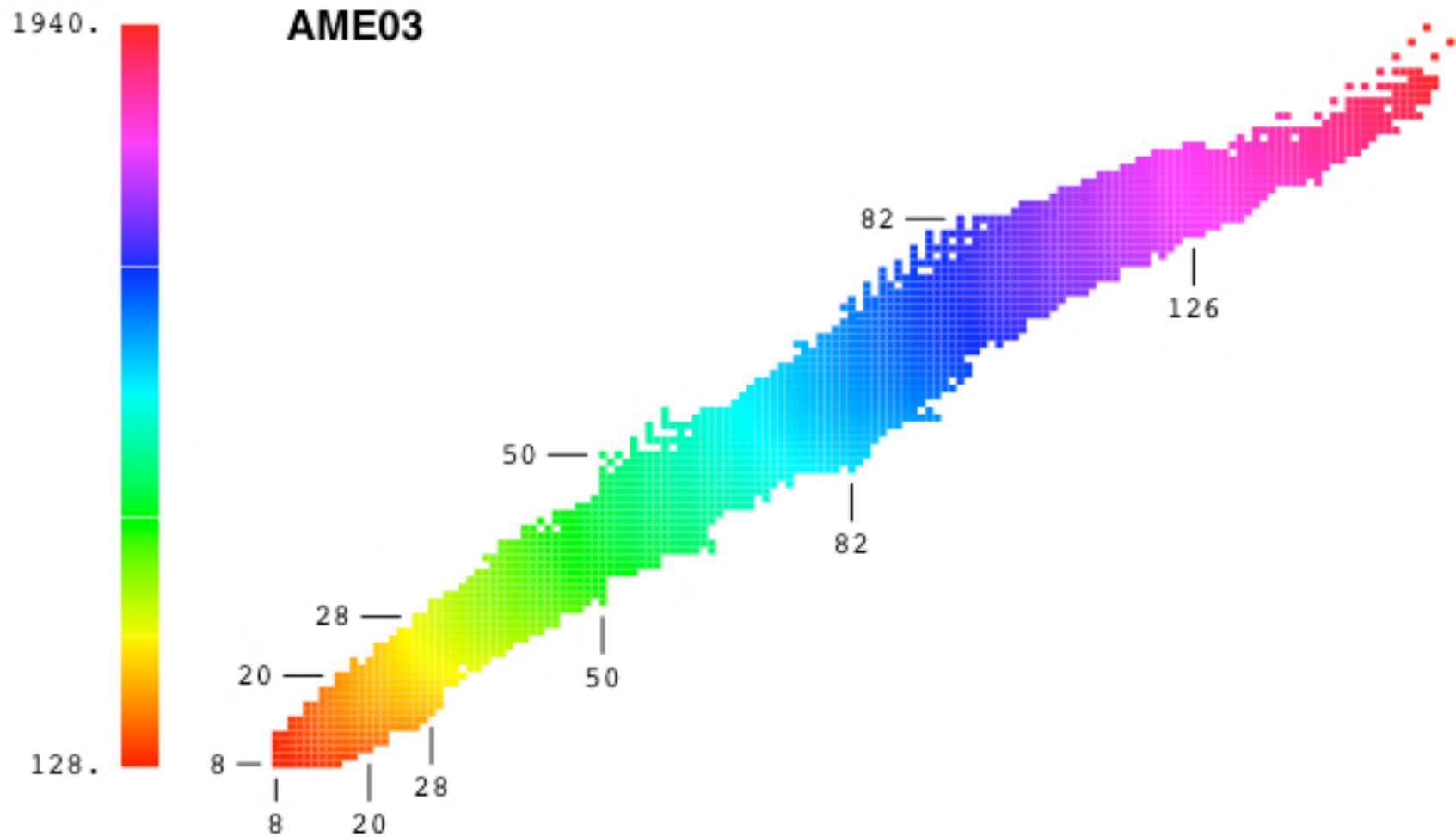
# Modified mass formula

- Add Wigner and surface-symmetry energy:

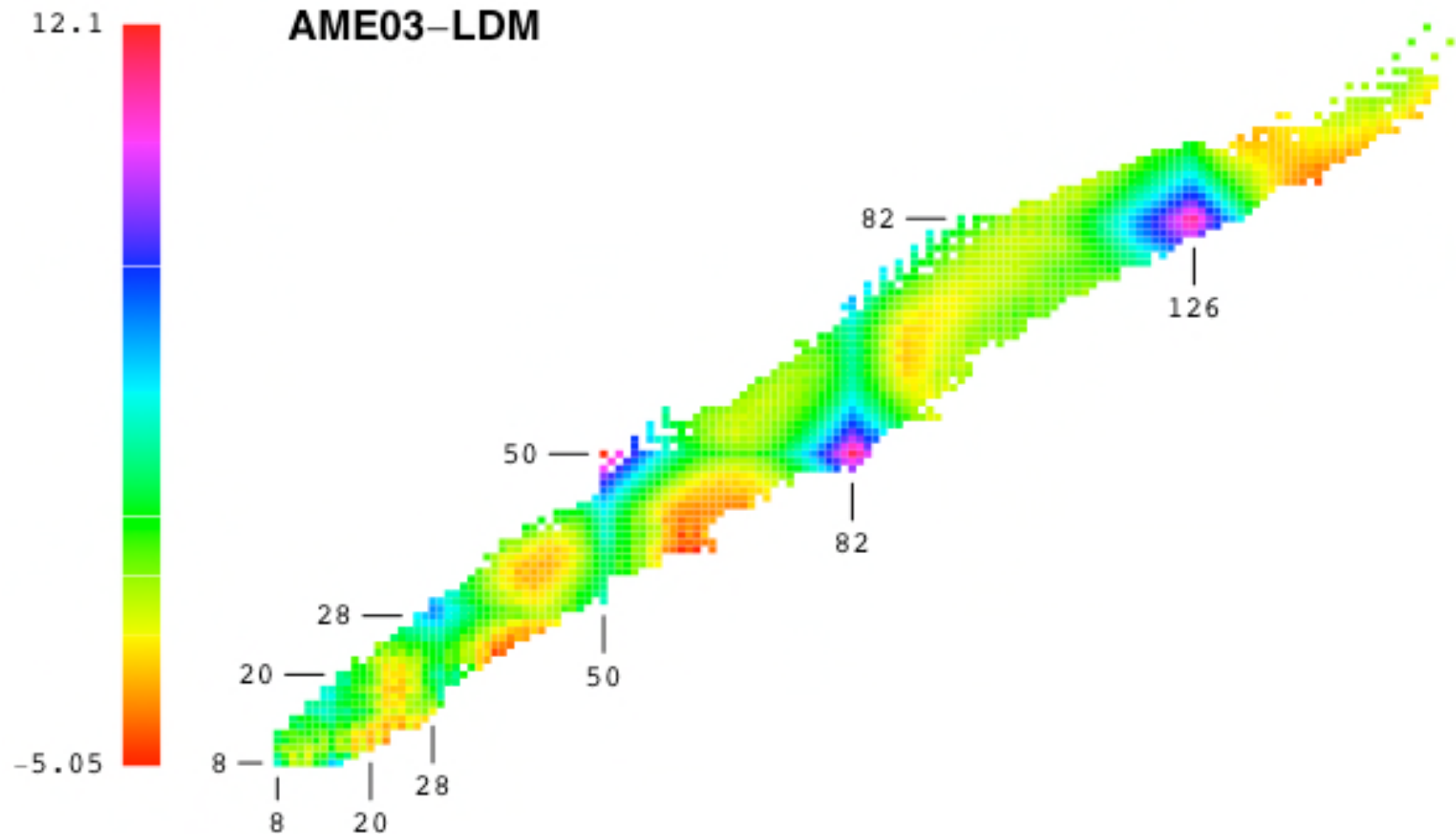
$$B(N, Z) = a_{\text{vol}} A - a_{\text{sur}} A^{2/3} - a_{\text{cou}} \frac{Z(Z-1)}{A^{1/3}} - \frac{S_v}{1 + S_v A^{-1/3} / S_s} \frac{T(T+r)}{A} + a_{\text{pai}} \frac{\delta(N, Z)}{A^{1/2}}$$

- Fit to AME03:  $\sigma_{\text{rms}} \approx 2.4 \text{ MeV}$ .

# The nuclear mass surface



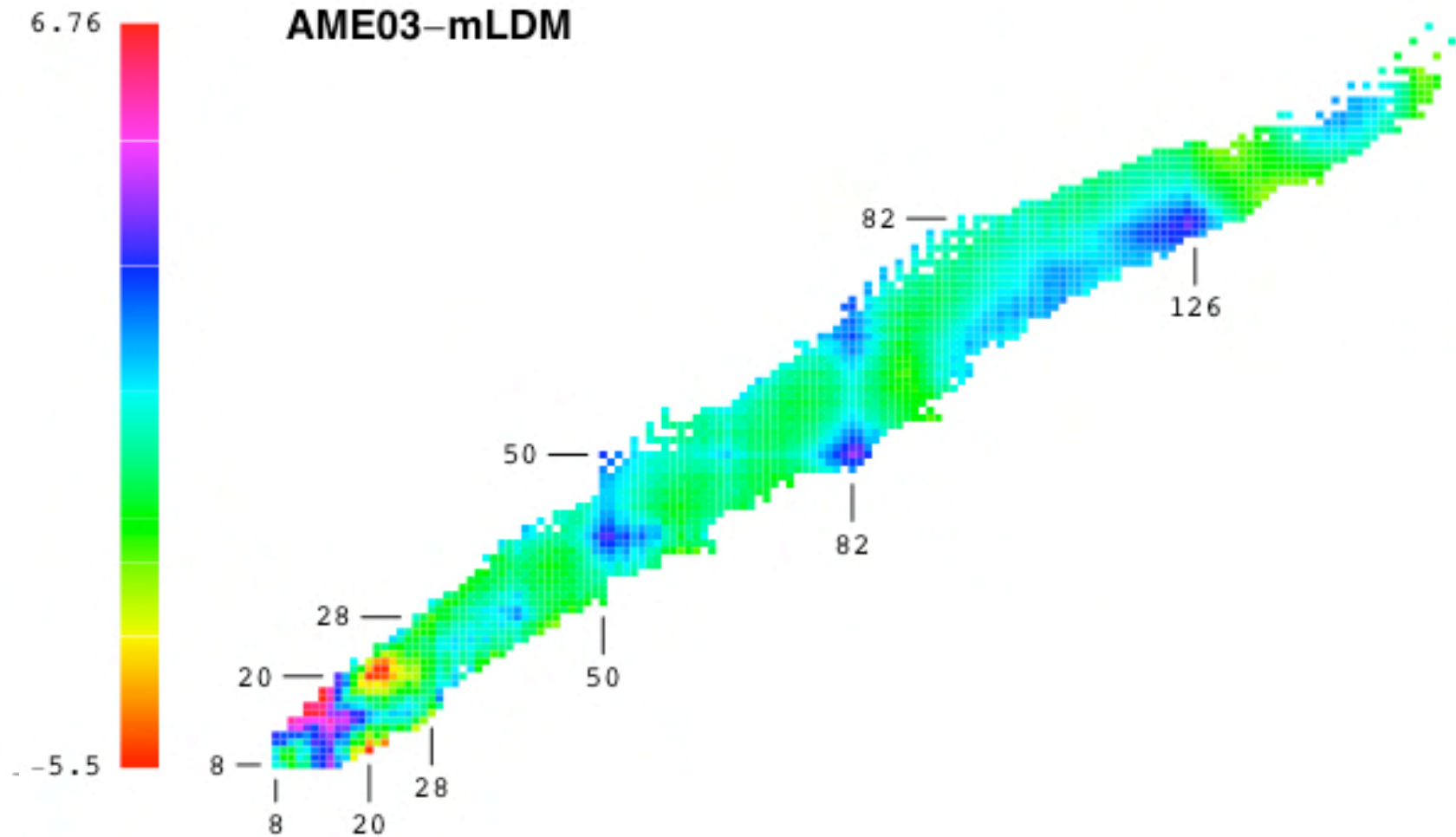
# The 'unfolding' of the mass surface



# Shell corrections

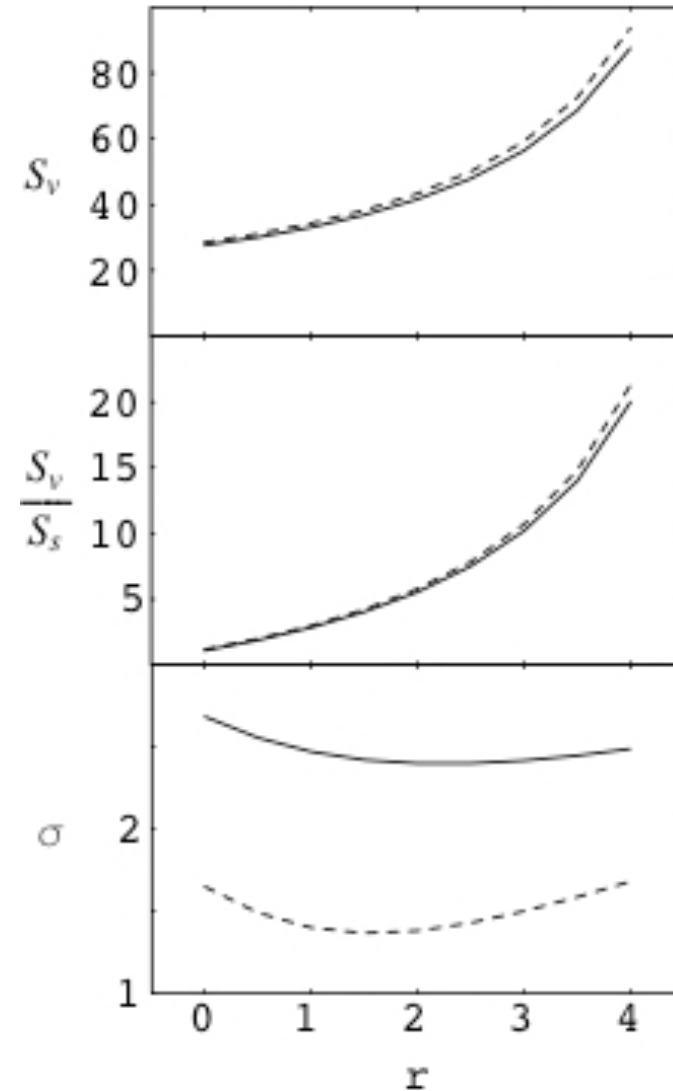
- Observed deviations suggest shell corrections depending on  $n_v+n_\pi$ , the total number of valence neutrons + protons (particles or holes).
- A simple parametrisation consists of two terms, linear and quadratic in  $F_{\max}=(n_v+n_\pi)/2$ .

# Shell-corrected LDM



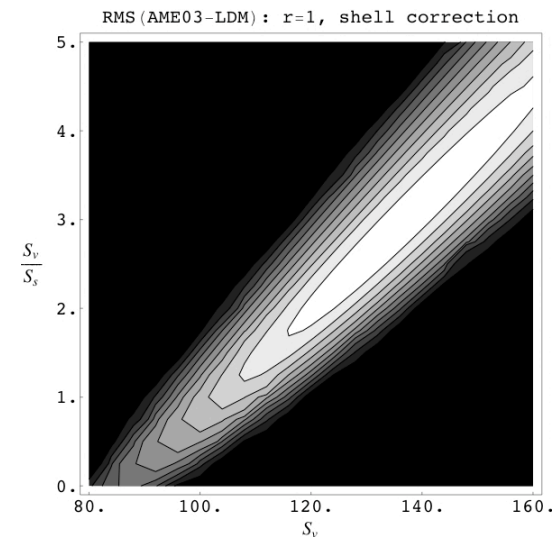
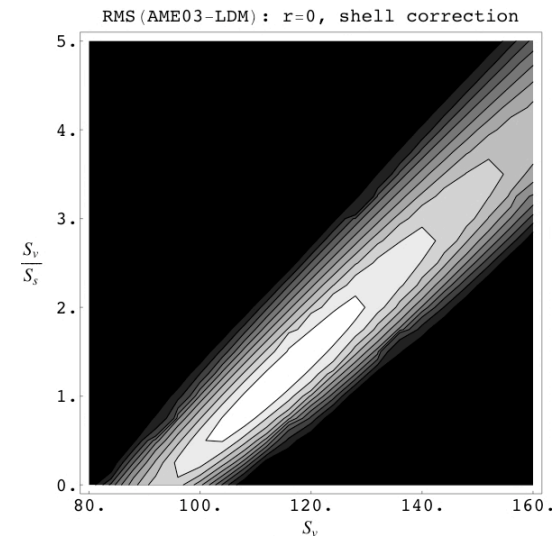
# Dependence on $r$

- Rms deviation  $\sigma$  decreases significantly with shell corrections.
- Rms deviation  $\sigma$  has shallow minimum in  $r$ .
- Both  $S_v$  and  $S_v/S_s$  are ill determined.



# Correlations

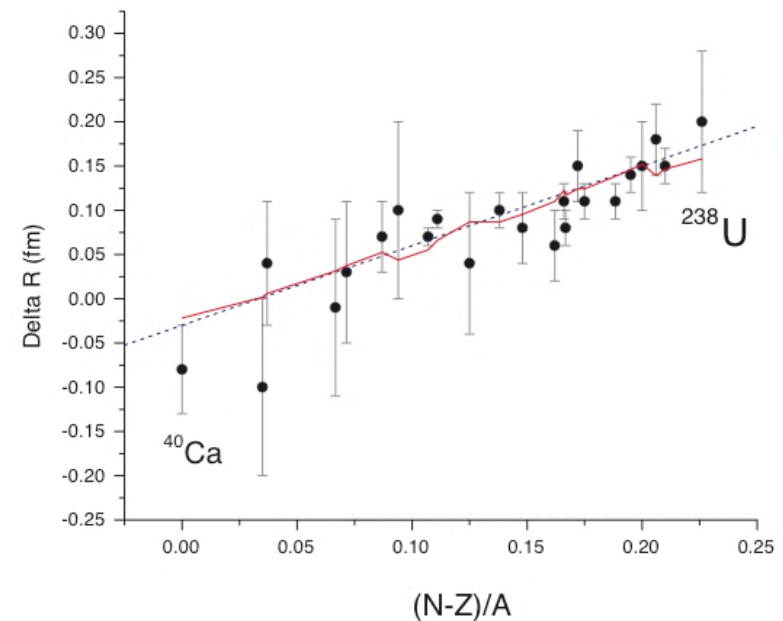
- Volume- and surface-symmetry terms are correlated.
- Correlation depends strongly on  $r$ .
- What nuclear properties are needed to determine  $S_v$  and  $S_s$ ?
- How can we fix  $r$ ?



# $S_V$ and $S_S$ from neutron skins

- Dependence of neutron skin on  $S_V$  and  $S_S$  (hard-sphere approximation):

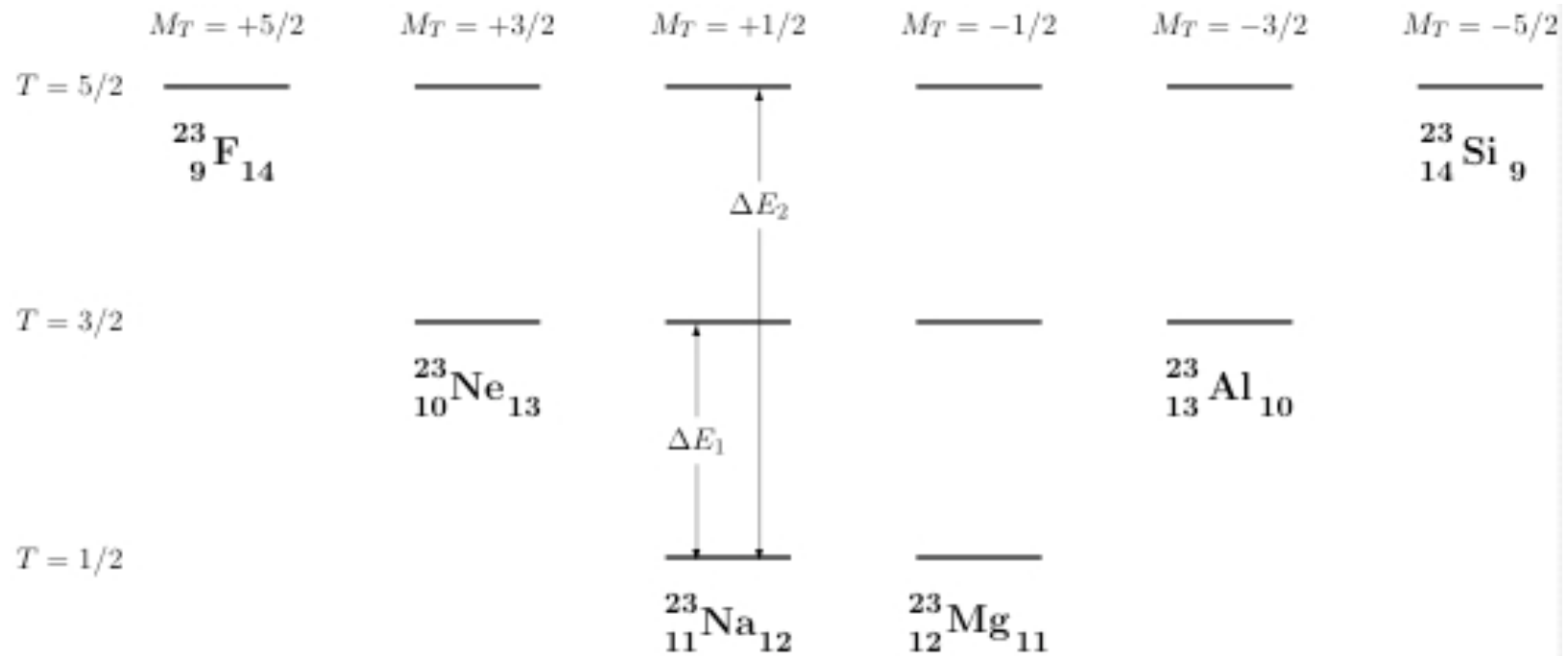
$$\frac{R_n - R_p}{R} = \frac{A}{6NZ} \frac{N - Z - a_{\text{cou}} A^{2/3} / (12S_V)}{1 + S_V A^{1/3} / S_S}$$



P. Danielewicz, Nucl. Phys. A **727** (2003) 233  
 A.E.L. Dieperink & P. Van Isacker, to be published



# $r$ from isobaric multiplets



- From  $T=3/2$  and  $T=5/2$  states in  $M_T = \pm 1/2$  nuclei:

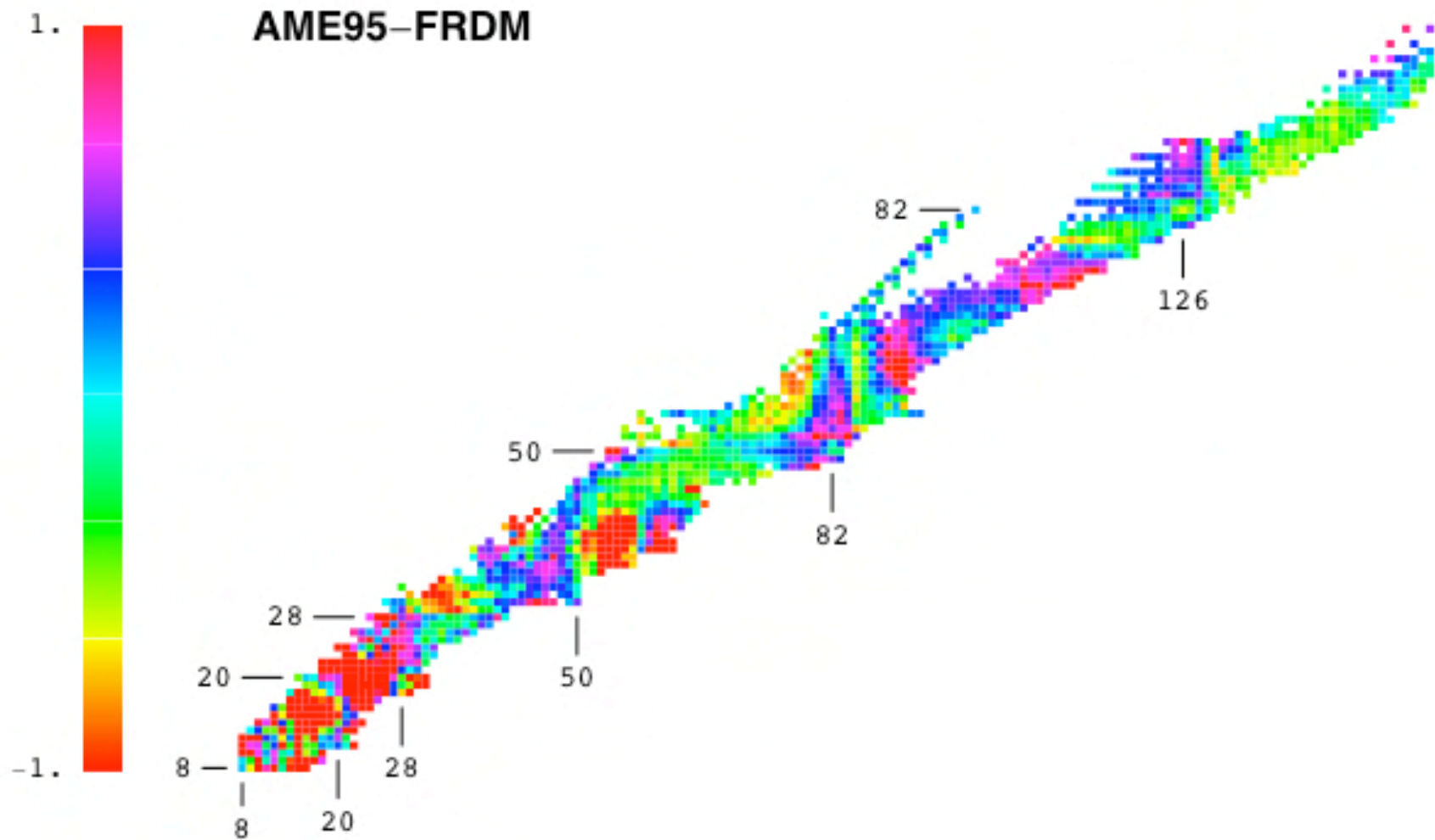
$$\frac{\Delta E_2}{\Delta E_1} = \frac{2(r+3)}{r+2} \Rightarrow r = \frac{6\Delta E_1 - 2\Delta E_2}{\Delta E_2 - 2\Delta E_1}$$

- In  ${}^{23}\text{Na}$ :  $\Delta E_1 = 7.891$  &  $\Delta E_2 = 19.586 \Rightarrow r = 2.15$

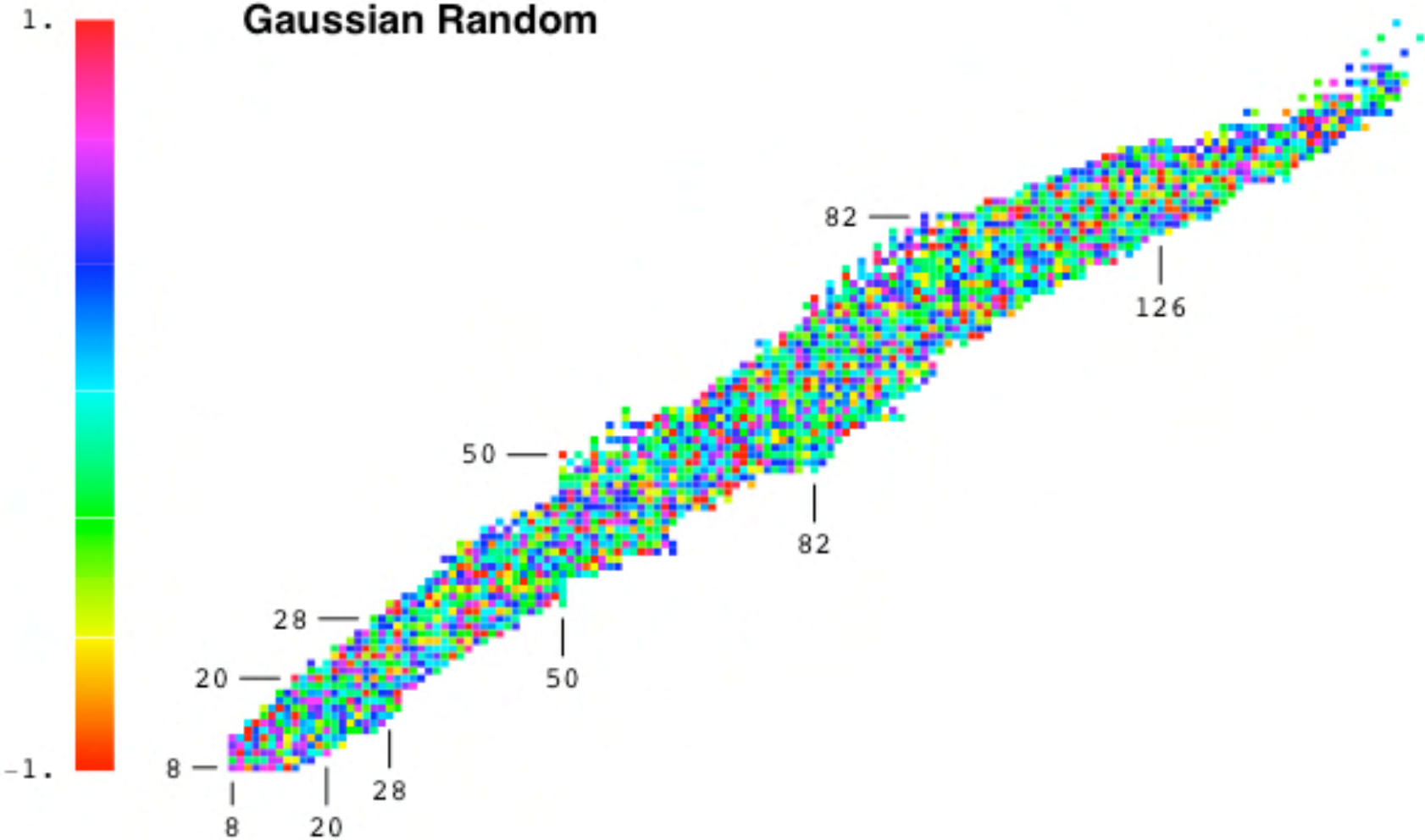
J. Jänecke & T.W. O'Donnell, Phys. Lett. B **605** (2005) 87

ALMAS-1, GSI-Darmstadt, October 2006

# Deviations from FRDM

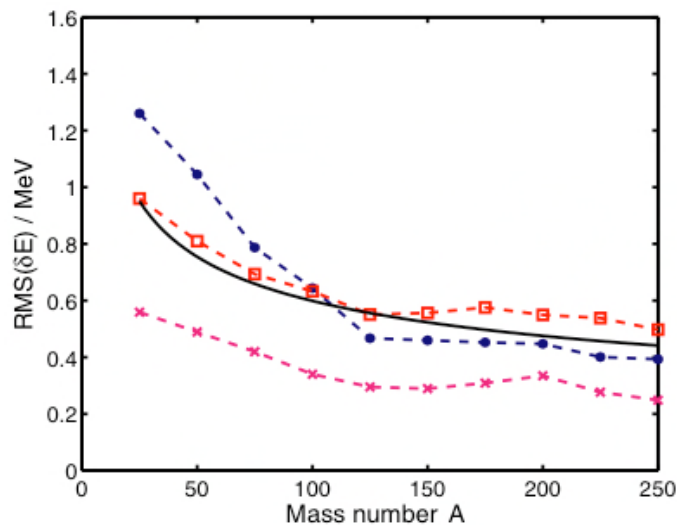


# Uncorrelated deviations



# Nuclear masses and chaos

- Precision of actual mass formulas is  $\sim 0.5$  MeV.
- Bohigas & Leboeuf: Current mass formulas assume regular dynamics and neglect a ‘chaotic’ deviation.
- Åberg: [...] it will certainly be extremely hard to improve the theoretical description of nuclear masses.



$$\sigma_{\text{chaos}}(N, Z) = \frac{2.78}{A^{1/3}} \text{ MeV}$$

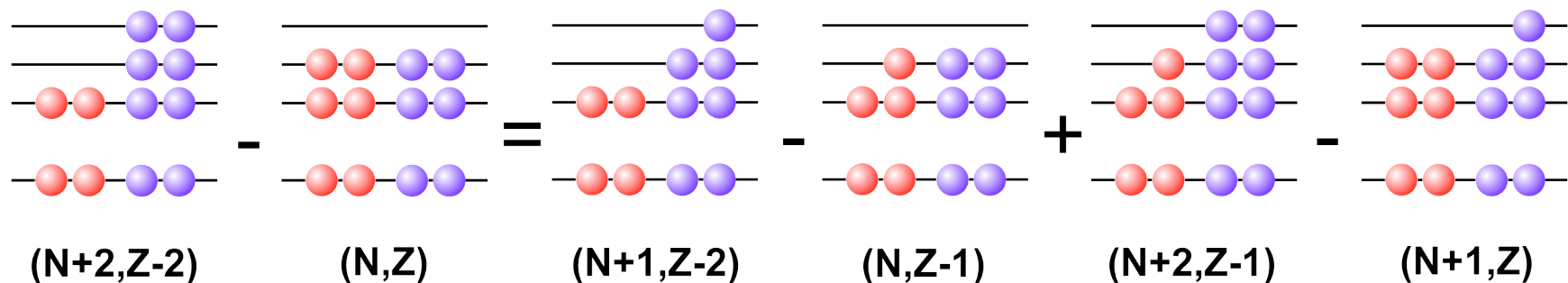
O. Bohigas & P. Leboeuf, Phys. Rev. Lett. **88** (2002) 092502  
S. Åberg, Nature **417** (2002) 499

# Garvey-Kelson relations

- Relations between masses of neighbouring nuclei:

$$-B(N+1, Z-2) + B(N+1, Z) - B(N+2, Z-1) + B(N+2, Z-2) - B(N, Z) + B(N, Z-1) = 0$$

$$B(N+2, Z) - B(N, Z-2) + B(N+1, Z-2) - B(N+2, Z-1) + B(N, Z-1) - B(N+1, Z) = 0$$



# Extended Garvey-Kelson relations

- The mass of every nucleus can be “predicted” on the basis of its neighbours in 12 different ways. The average gives:

	$N - 2$	$N - 1$	$N$	$N + 1$	$N + 2$
$Z - 2$	+1	-2	+2	-2	+1
$Z - 1$	-2		+4		-2
$Z$	+2	+4	-12	+4	+2
$Z + 1$	-2		+4		-2
$Z + 2$	+1	-2	+2	-2	+1

# Mass deviations and errors

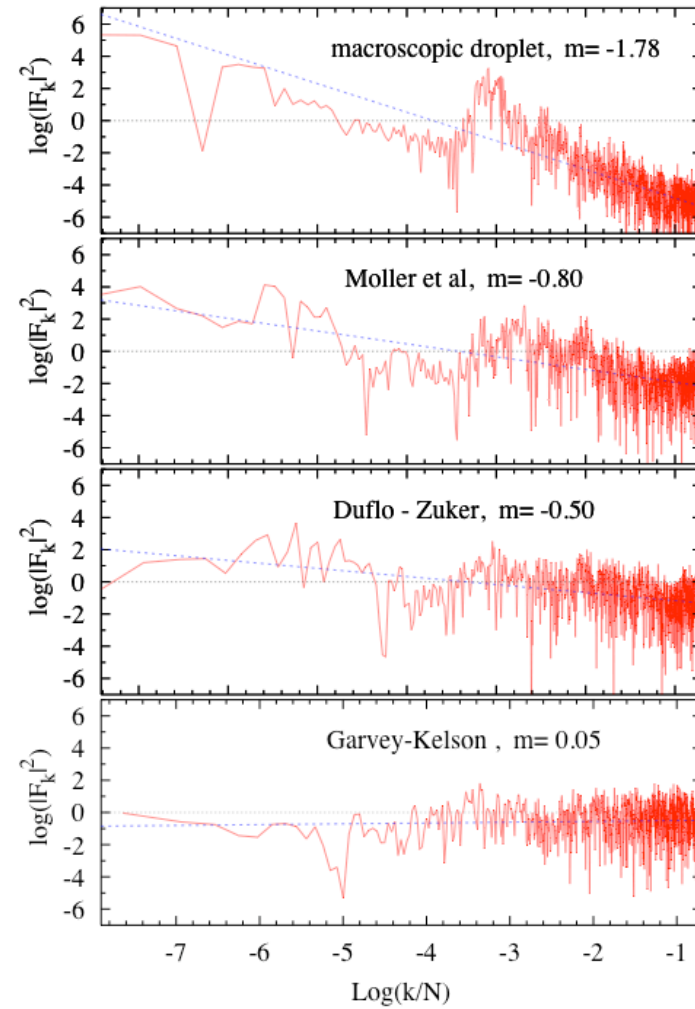
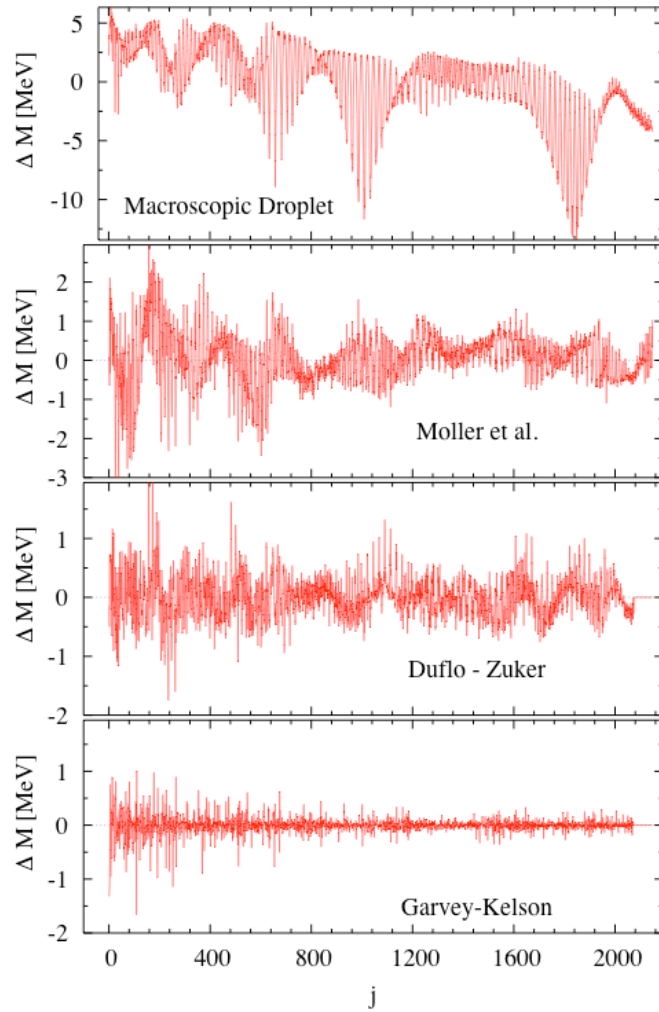
- Deviations  $\sigma_{\text{rms}}$  and errors  $\sigma_{\text{err}}$ :

$$\sigma_{\text{rms}} = \sqrt{\frac{1}{M} \sum_{j=1}^M (B_j^{\text{expt}} - B_j^{\text{theo}})^2}, \quad \sigma_{\text{err}} = \sqrt{\frac{1}{M} \sum_{j=1}^M (\Delta B_j^{\text{expt}})^2}$$

- Deviation and errors (in MeV) for  $N, Z \geq 8$ :

	mLDM	FRDM	HFB9	DZ	GK1	GK12
$M$	2149	2149	2149	2149	2066	1008
$\sigma_{\text{rms}}$	1.187	0.650	0.733	0.360	0.155	0.087
$\sigma_{\text{err}}$	0.121	0.121	0.121	0.121	0.111	0.034

# LDM, FRDM, DZ, GK fluctuations





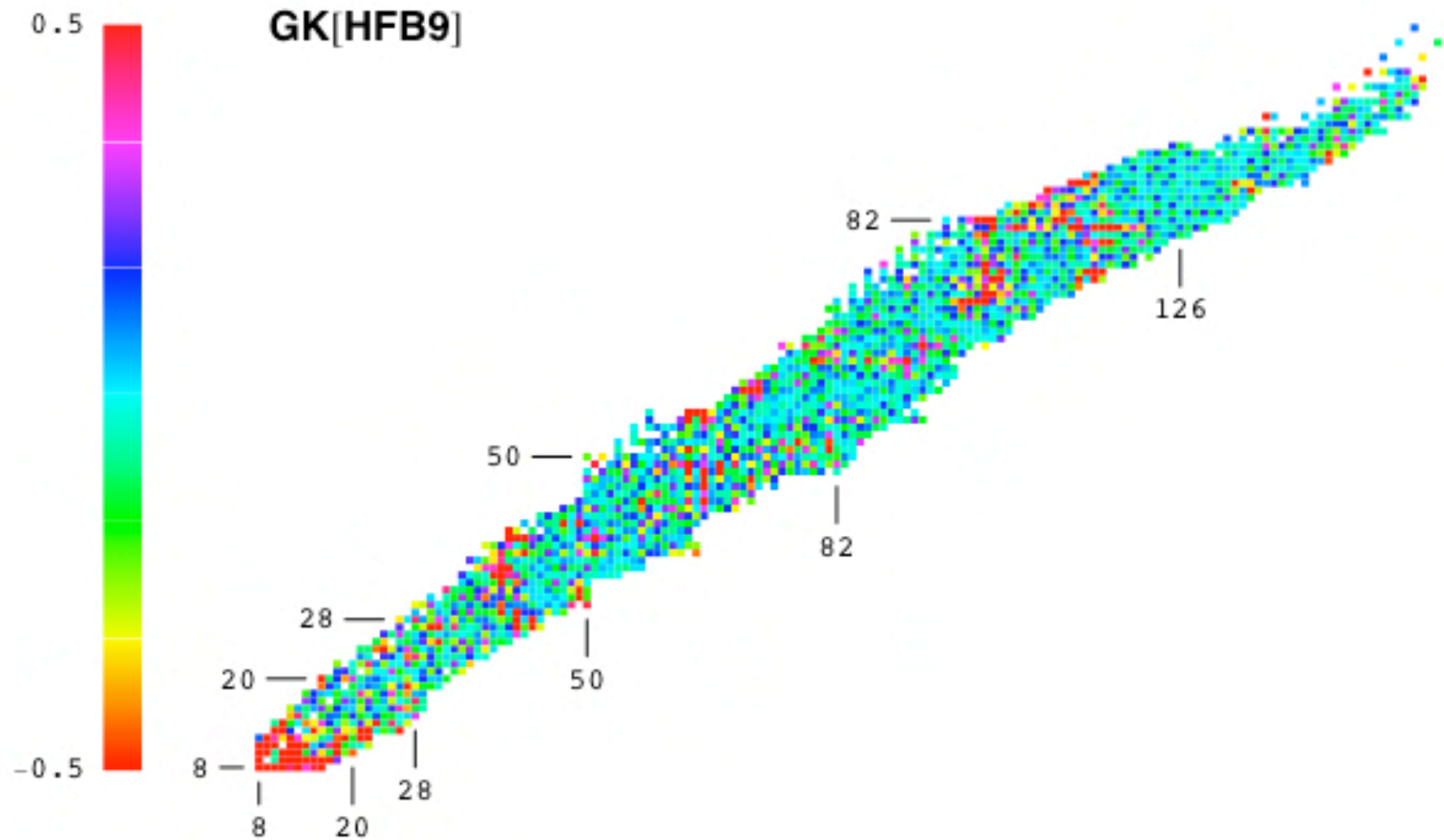
# Correlated chaos

- The Garvey-Kelson relations provide an example of uncorrelated deviations (white noise) with  $\sigma_{\text{rms}} \approx 100 \text{ keV}$ .
- The chaotic mass component is not entirely uncorrelated and is, at least partially, predictable.

*“It’s chaos but organized chaos”*

Charles Mingus

# Tests of mass predictions from GK



# Test of mass predictions from GK

- Do mass predictions satisfy the GK relations?  
Example: How well is GK12 satisfied?

	All in <i>M</i>	nuclei AME95 $\sigma_{\text{rms}}$	" All" nuclei <i>M</i>	$\sigma_{\text{rms}}$	" All" not in <i>M</i>	nuclei AME95 $\sigma_{\text{rms}}$
Expt	570	0.098	--	--	--	--
FRDM	1555	0.125	7223	0.280	5668	0.309
HFB9	1555	0.212	6286	0.339	4731	0.372
DZ	1555	0.037	7188	0.045	5633	0.047

# Test of mass predictions from GK

- Do mass predictions satisfy the GK relations?  
Example: How well is GK12 satisfied?

	All in	nuclei AME03	" All" nuclei	" All" not in	nuclei AME03	
	$M$	$\sigma_{\text{rms}}$	$M$	$\sigma_{\text{rms}}$	$M$	$\sigma_{\text{rms}}$
Expt	1008	0.087	--	--	--	--
FRDM	1909	0.126	7223	0.280	5314	0.318
HFB9	1918	0.226	6286	0.339	4368	0.379
DZ	1918	0.037	7188	0.045	5270	0.047

# Conclusions

- Consider surface *and* Wigner corrections in the liquid-drop mass formula to determine the symmetry energy in nuclei.
- Chaotic mass component is *not* unpredictable and can be (at least partially) calculated.
- Measured masses do satisfy GK relations; several existing mass formulas do not.